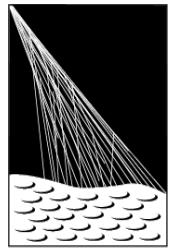


PIERRE
AUGER
OBSERVATORY

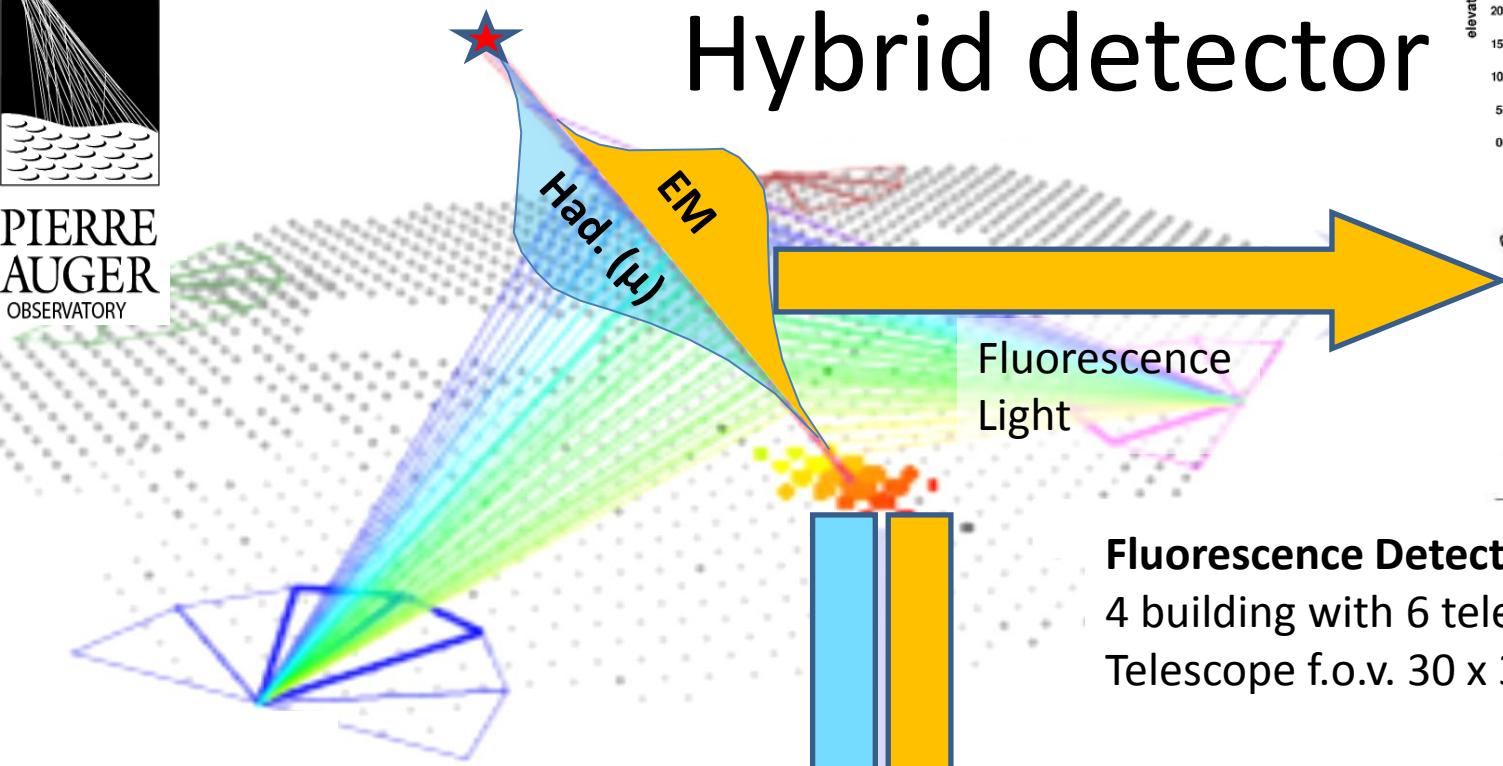


Constraints and measurements of hadronic interactions in extensive air showers with the Pierre Auger Observatory

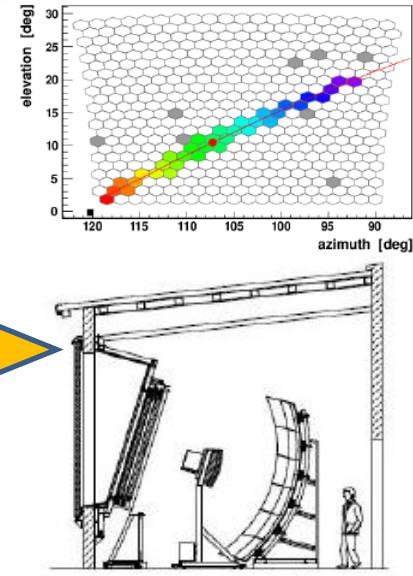
L. Cazon, for the Auger Collaboration



PIERRE
AUGER
OBSERVATORY



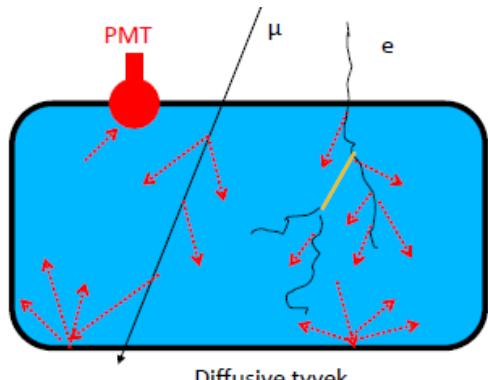
Hybrid detector



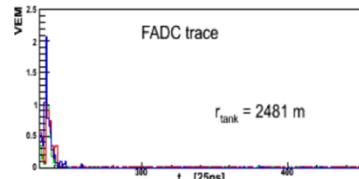
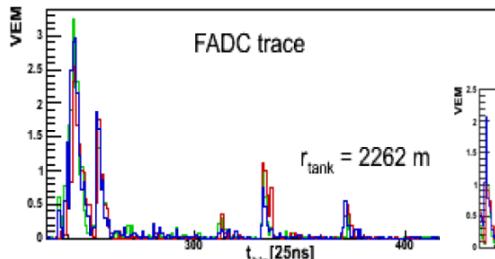
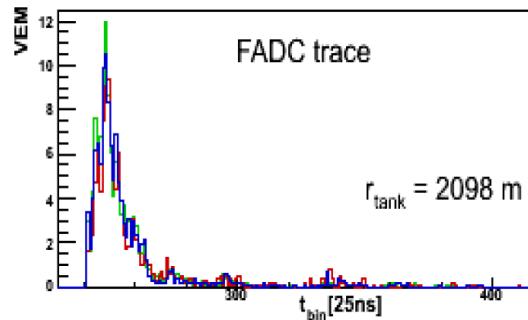
Fluorescence Detector (FD)
4 building with 6 telescopes each
Telescope f.o.v. 30×30 deg

Surface Detector (SD)

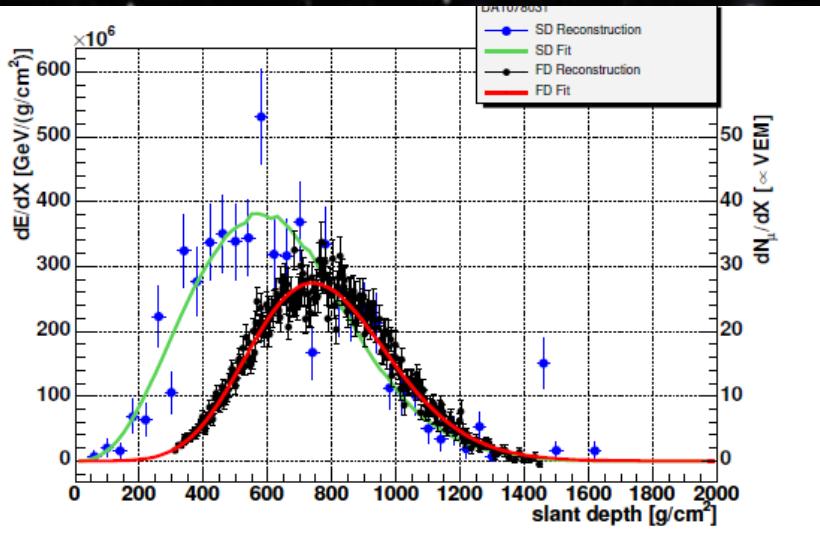
1600 water Cherenkov tanks
Area of 3000 km^2



SD & FD
Enhancements

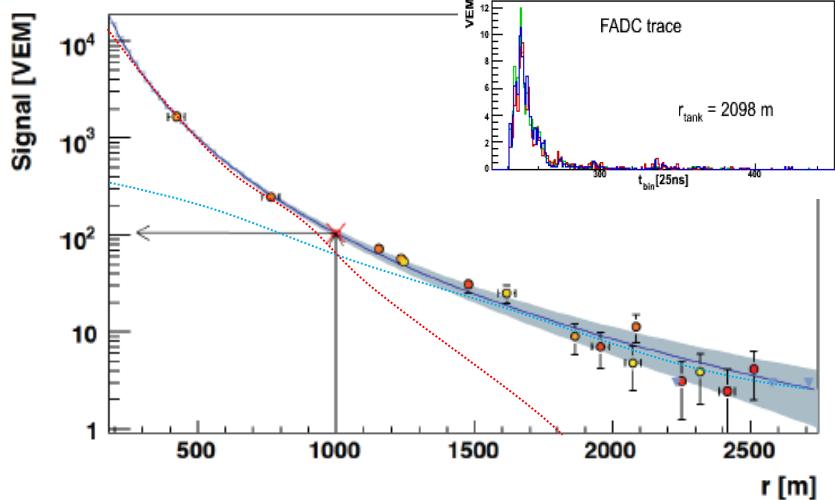


Auger observables at a glance



Longitudinal development

- EM profile (FD)
- Muon Production profile (SD)
 - Time structure of muon dominated regions



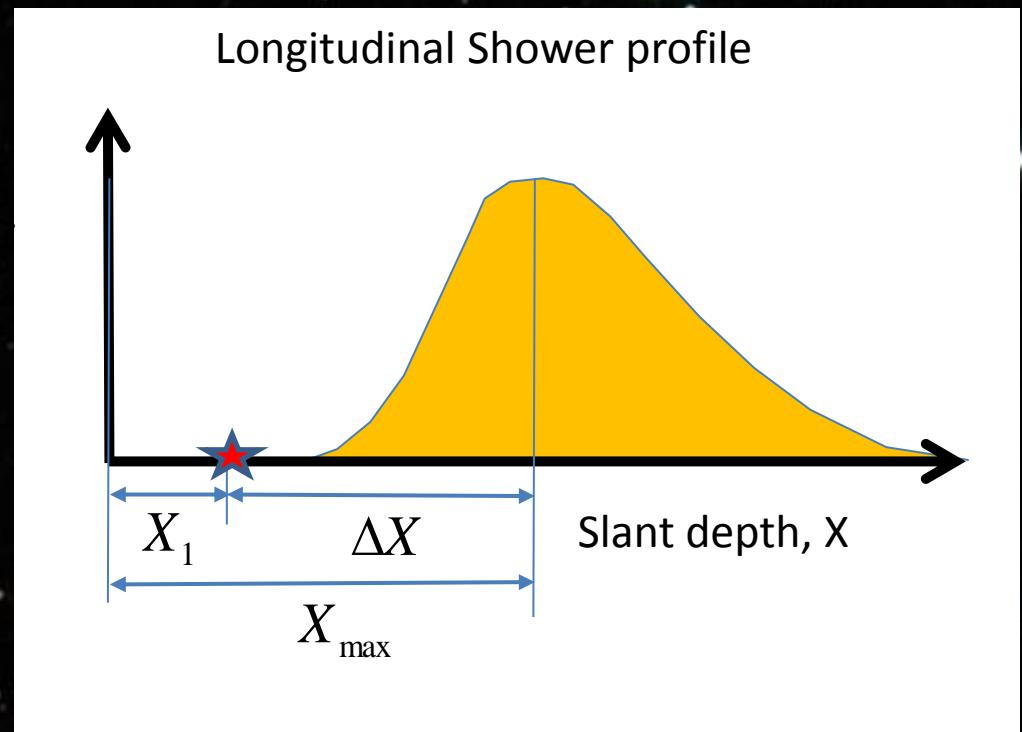
Lateral ground distribution

- All particles (total signal)
- Pure muonic
 - Inclined showers
- EM/Mu separation
 - Time structure

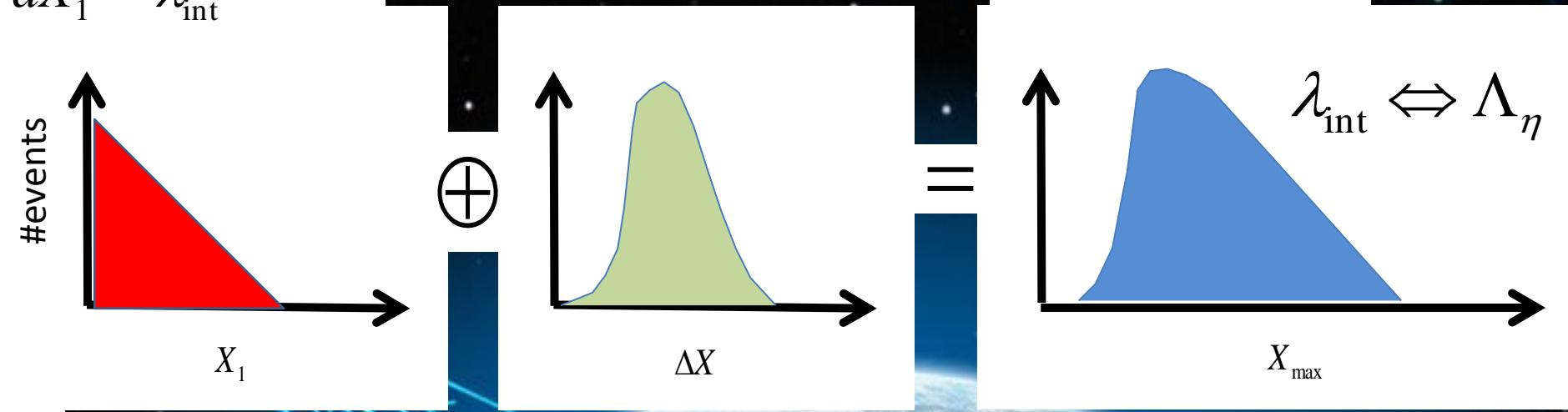
p-Air cross section

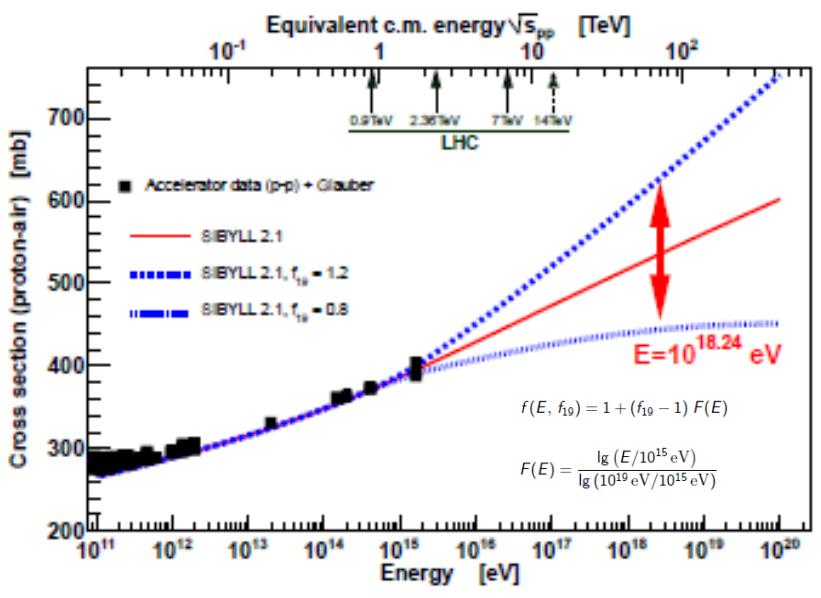
$$\sigma_{\text{int}} = \frac{\langle m_{\text{air}} \rangle}{\lambda_{\text{int}}}$$

$$\frac{dp}{dX_1} = \frac{1}{\lambda_{\text{int}}} e^{-X_1/\lambda_{\text{int}}}$$



⇒ Tail of X_{max} -Distribution
 $dN/dX_{\text{max}} \propto \exp(-X_{\text{max}}/\Lambda_\eta)$



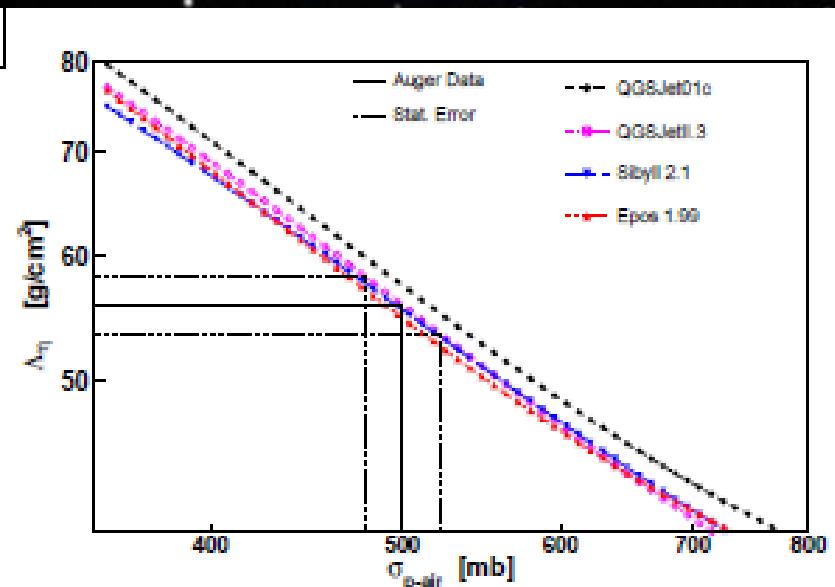
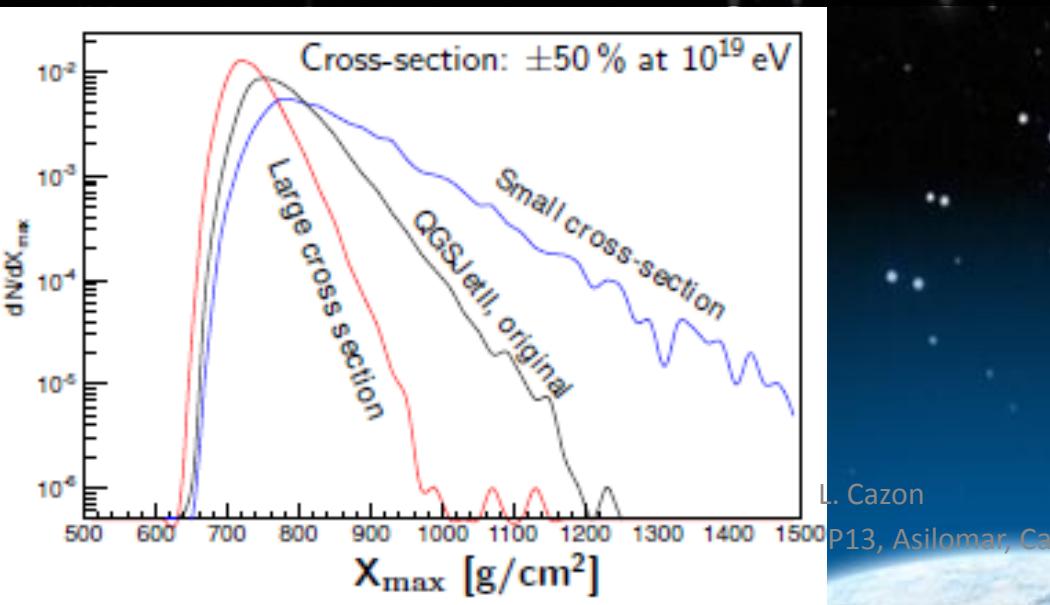


Method

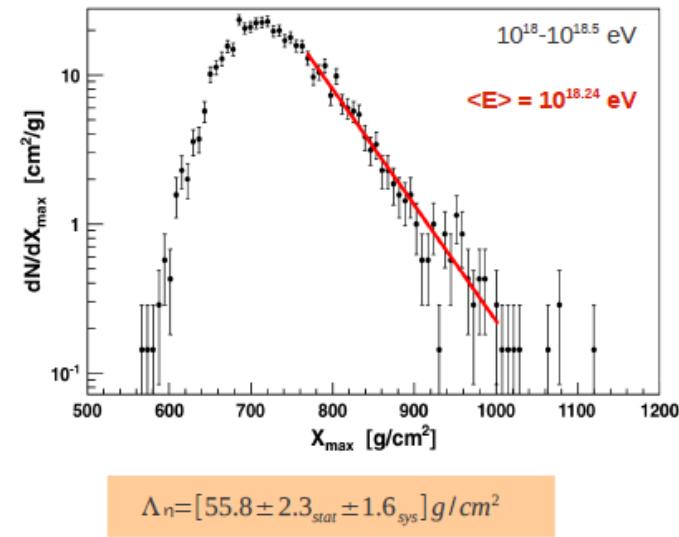
- Continuous reparameterization of cross section in MC

- Simulation of X_{\max} distribution
 - different rescalings
 - different models

- $\Lambda_\eta \leftrightarrow \sigma_{p\text{-Air}}$ conversion

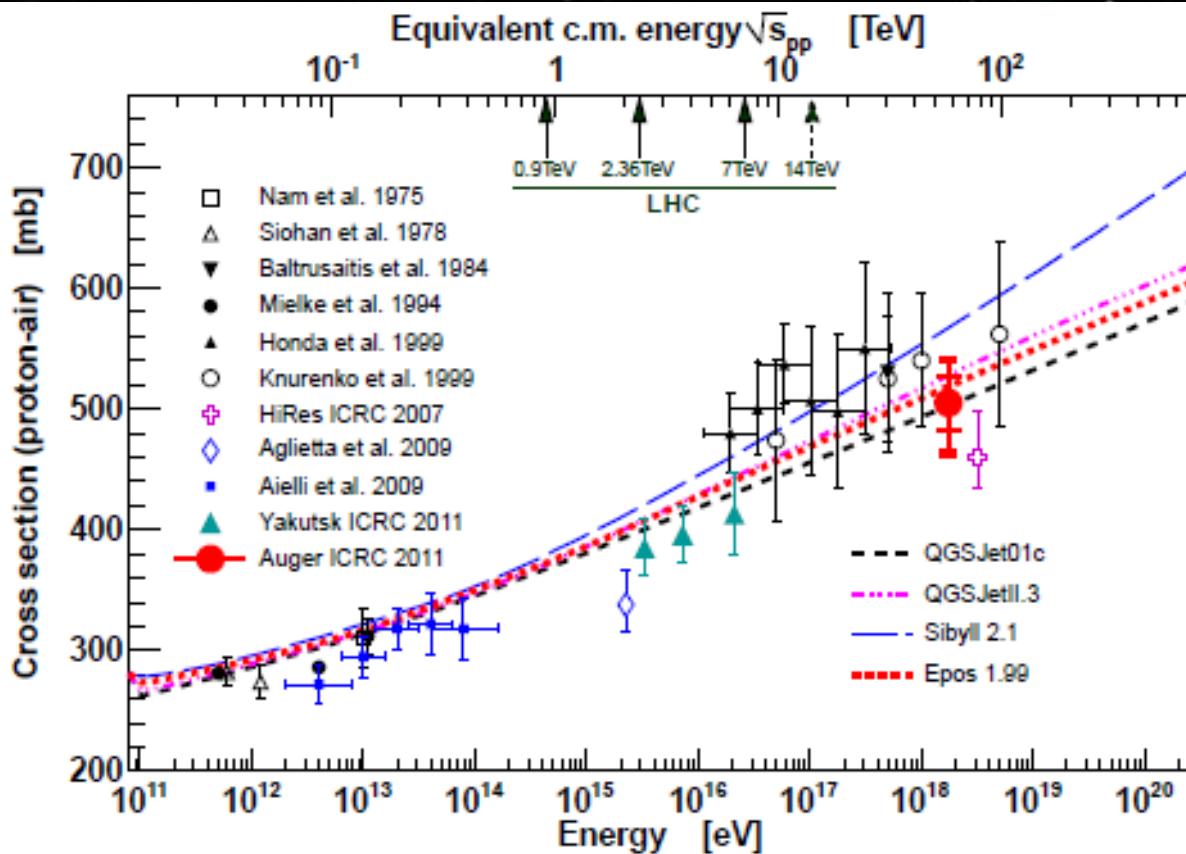


Results



Description	Impact on $\sigma_{p\text{-air}}$
Λ_η systematics	$\pm 15 \text{ mb}$
Hadronic interaction models	$\pm 19 \text{ mb}$
Energy scale	$\pm 7 \text{ mb}$
Conversion of Λ_η to $\sigma_{p\text{-air}}^{\text{prod}}$	$\pm 7 \text{ mb}$
Photons, <0.5 %	< +10 mb
Helium, 10 %	-12 mb
Helium, 25 %	-30 mb
Helium, 50 %	-80 mb
Total (25 % helium)	-36 mb, +28 mb

Possible He contamination is the main source of systematic uncertainty. 25% He maximum contamination assumed for sys. uncertainties



$$\sigma_{p\text{-air}} = [505 \pm 22_{\text{stat}} \quad (+28)_{\text{sys}} \quad (-36)] \text{ mb}$$

$$\sigma_{pp}^{\text{inel}} = [90 \pm 7_{\text{stat}} \quad (+9)_{\text{sys}} \quad \pm 1.5_{\text{Glauber}}] \text{ mb}$$

$$\sqrt{s_{pp}} = [57 \pm 6] \text{ TeV}$$

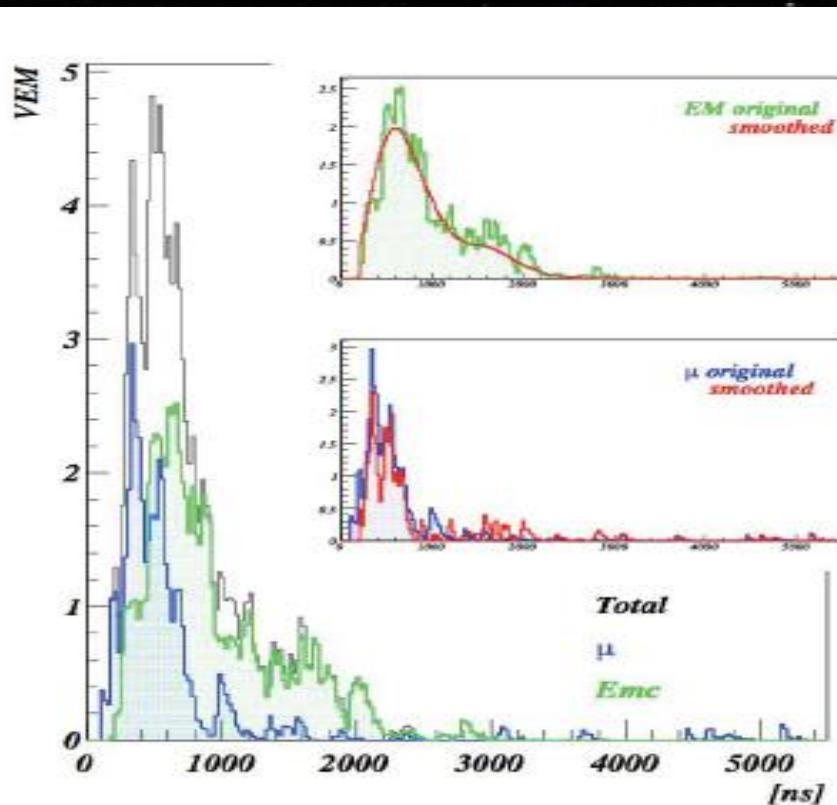
Muon production in air showers

- Analysis of temporal structure of SD signals
- Analysis of signal size of inclined showers
- Analysis of signal size of hybrid events

Temporal structure of SD signals

1. Smoothing

Low pass filter-> EM component
After Subtraction of EM-> Muons



2. Multivariate Analysis

Variables sensitive to large relative fluctuations and short signals.

$$f_{0.5}^2 = \frac{\left(\sum x_j (x_j > 0.5) \right)^2}{\left(\sum x_j \right)^2} \quad P_0 = \frac{\langle x \rangle^2}{\langle x^2 \rangle}$$

Conversion to muon fraction by:

$$f_\mu = a + b\theta + c f_{0.5}^2 + d\theta P_0 + e r$$

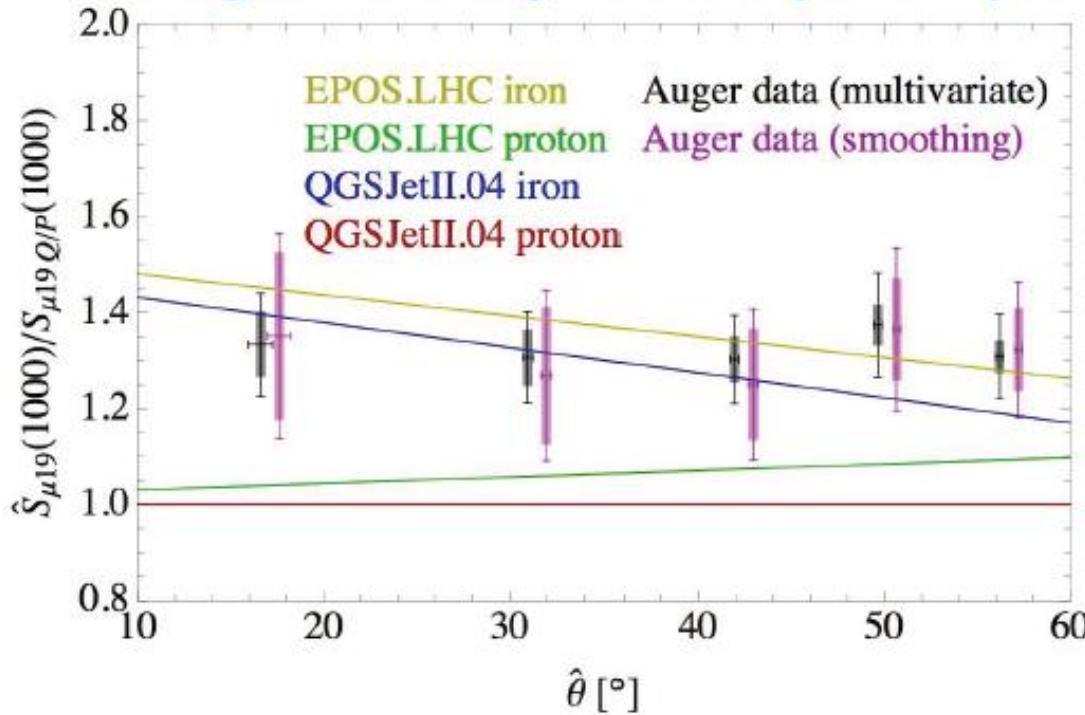
Where a, b, c, d, e were calculated using simulations

See more details ICRC2013

Results

$E = [10^{18.98}, 10^{19.02}] \text{ eV}$
 $r = [950, 1050] \text{ m}$

Muon signal rescaling wrt QGSJetII.04 proton

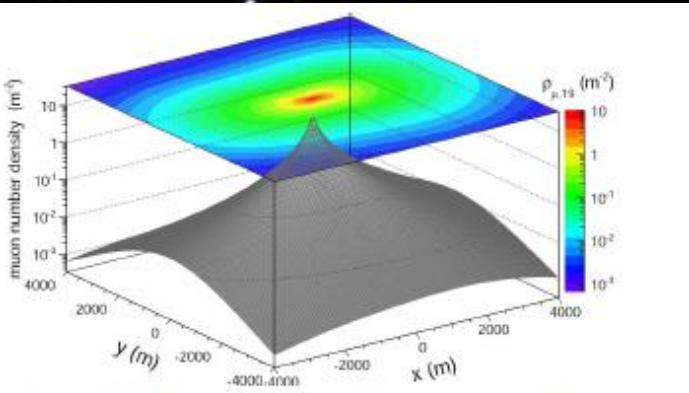
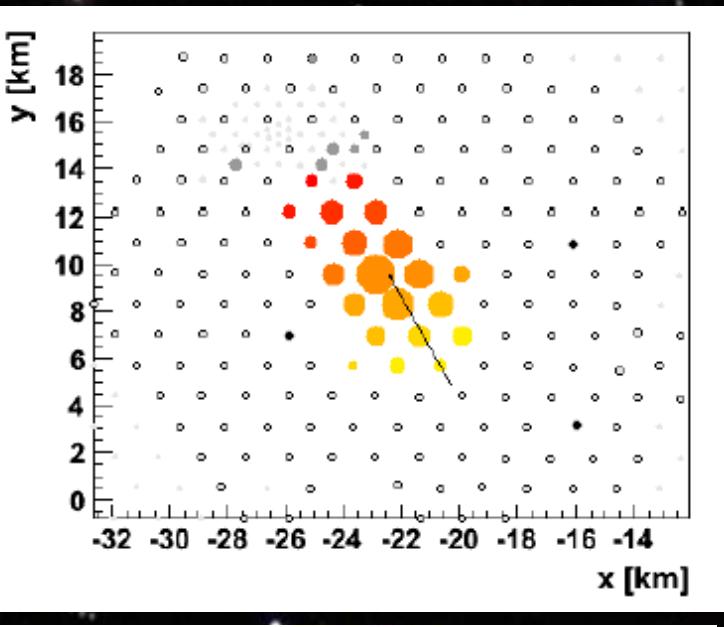


$1.33 \pm 0.02 \text{ (stat.)} \pm 0.05 \text{ (sys.)} \text{ (multivariate)}$

$1.31 \pm 0.02 \text{ (stat.)} \pm 0.09 \text{ (sys.)} \text{ (smoothing)}$

Inclined hybrid events

$62 < \theta < 80$ deg



Example of $\rho_{\mu,19}$ for proton showers at $\theta=80^\circ$, $\phi=0^\circ$ and core at $(x,y)=(0,0)$

Fit the muon density in stations

$$\rho_\mu = N_{19} \rho_{\mu,19}(x, y)$$

where N_{19} free parameter

And $\rho_{\mu,19}(x, y)$ is fixed, corresponding to proton QGSJetII-03 at 10^{19} eV

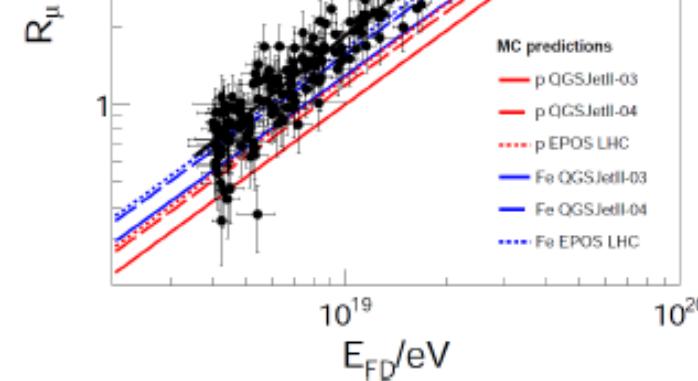
Ratio of the total number of muons N_μ to $N_{\mu,19}$ (proton QGSJetII-03 at 10^{19} eV)

$$R_\mu = N_\mu / N_{\mu,19}$$

Correspondence (<5% bias correction)

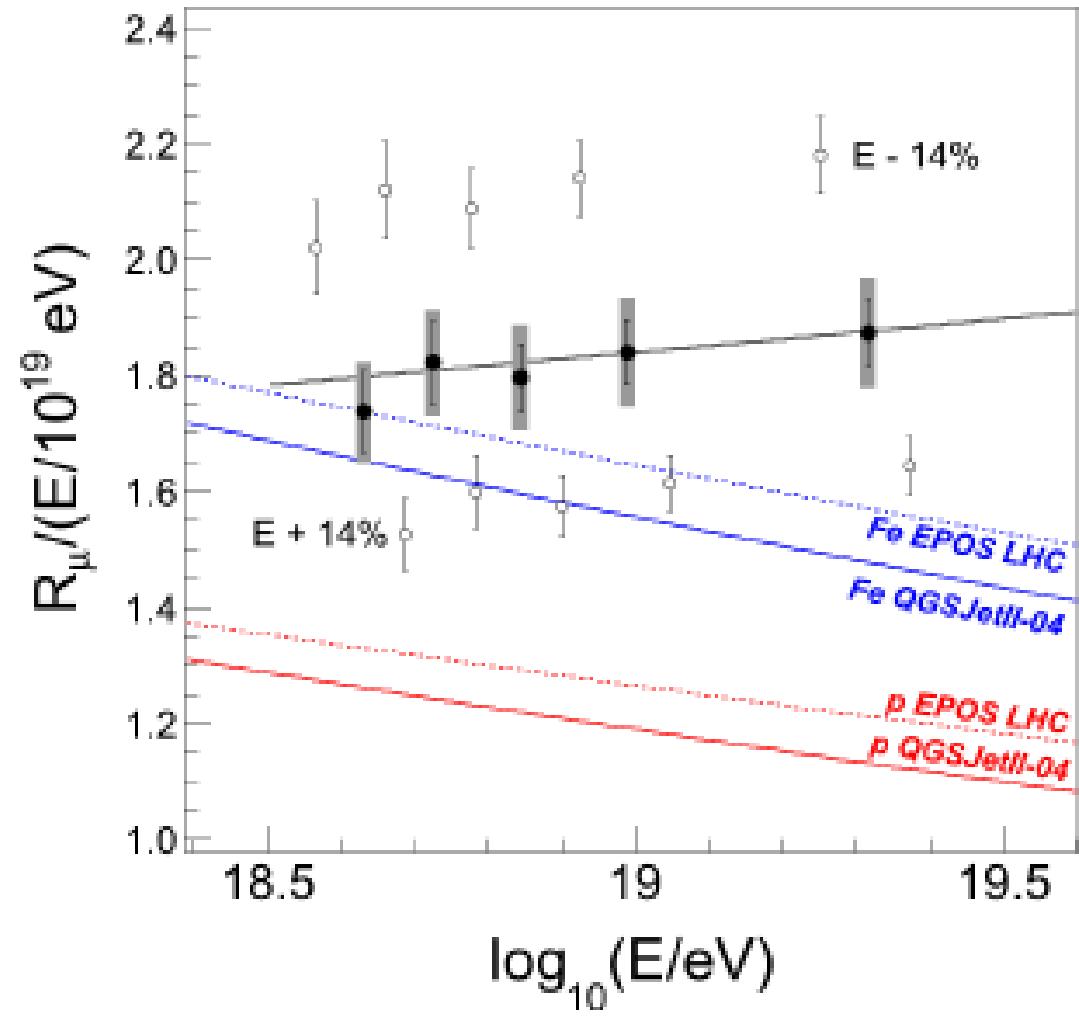
$$N_{19} \Leftrightarrow R_\mu$$

R_μ vs E_{FD}
(*calorimetric energy measured with the FD)

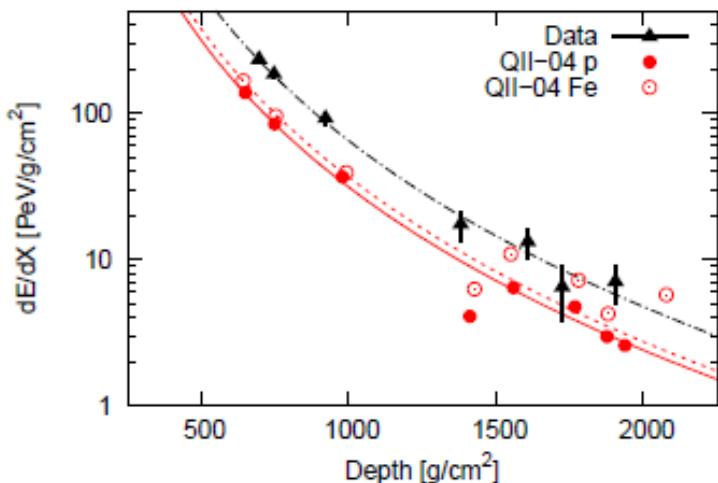
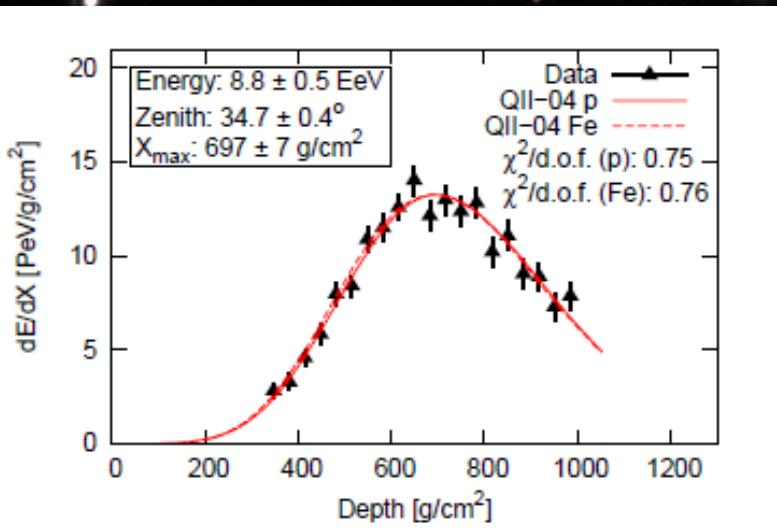


Results

R_μ/E_{FD} in energy bins



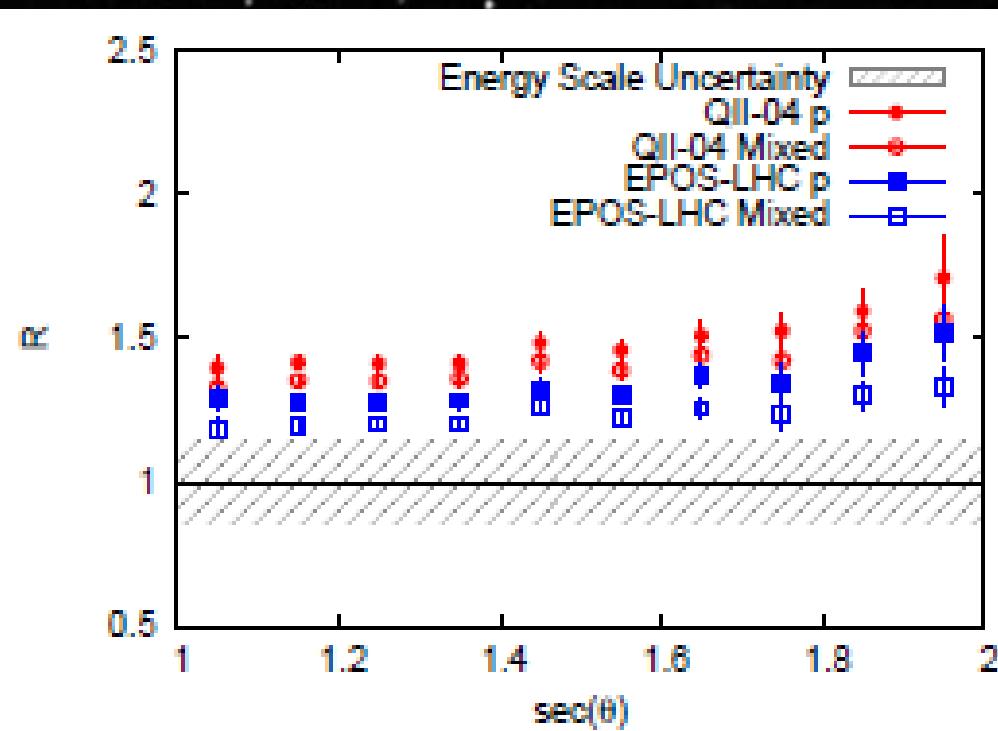
Hybrid events



$$E = [10^{18.8}, 10^{19.2}] \text{ eV}$$

- Find simulations which match FD profile, **for each event**
- Compare SD signals for simulations and data
- Rescale muon content until simulated SD best match data

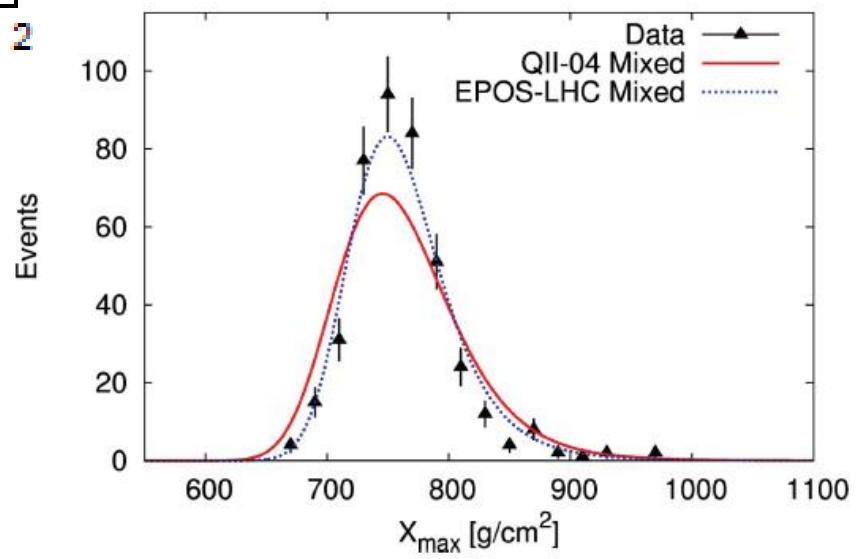
Results



L. Cazon

TAUP13, Asi

Mixed composition
reproduces X_{\max} -
distribution of data



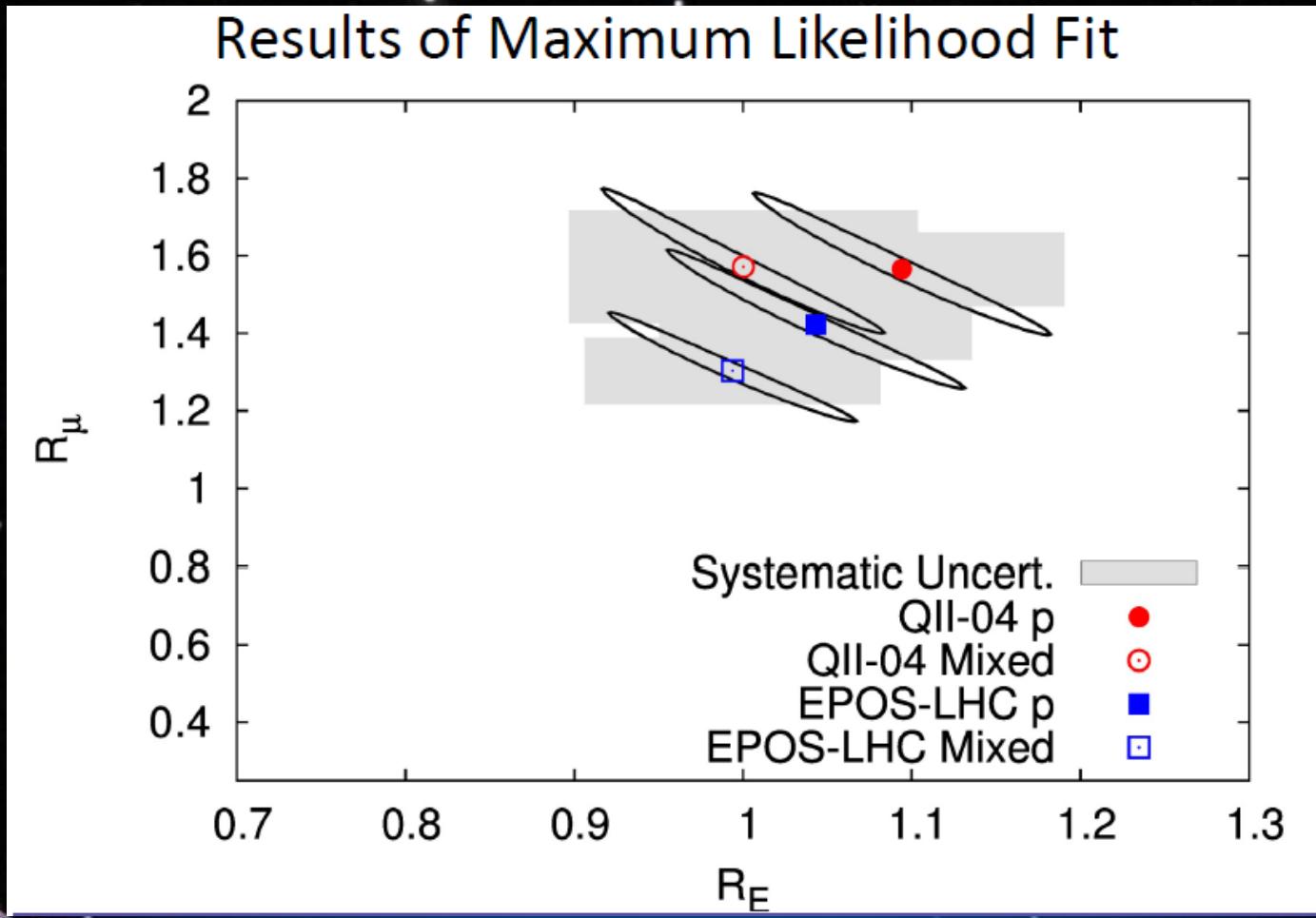
Identifying the discrepancy

- R_E : Energy rescaling. Rescales EM and muonic components
- R_μ : Muonic rescaling.
- Find R_E & R_μ for best overall fit

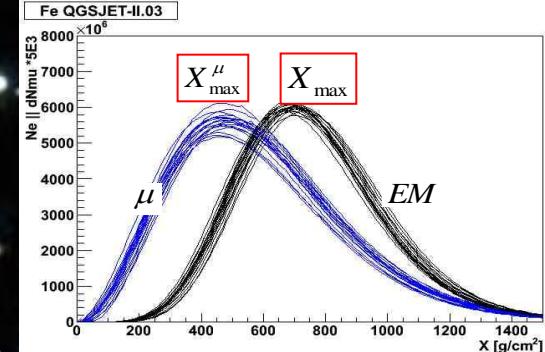
$$S_{resc}(R_E, R_\mu)_{i,j} \equiv R_E S_{EM,i,j} + R_E^\alpha R_\mu S_{\mu,i,j}$$

$$Likelihood = \prod_i \sum_j p_j(X_{\max,i}) Gaus(S_{resc}(R_E, R_\mu)_{i,j} - S_{1000i})$$

Results



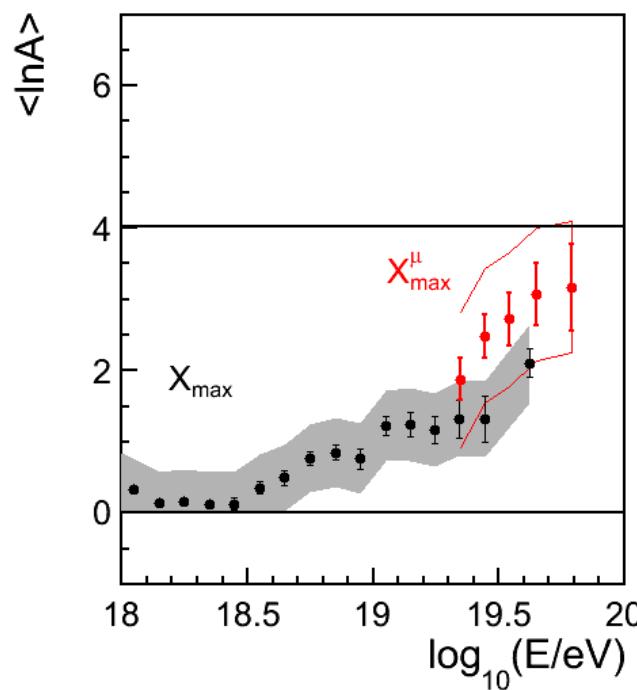
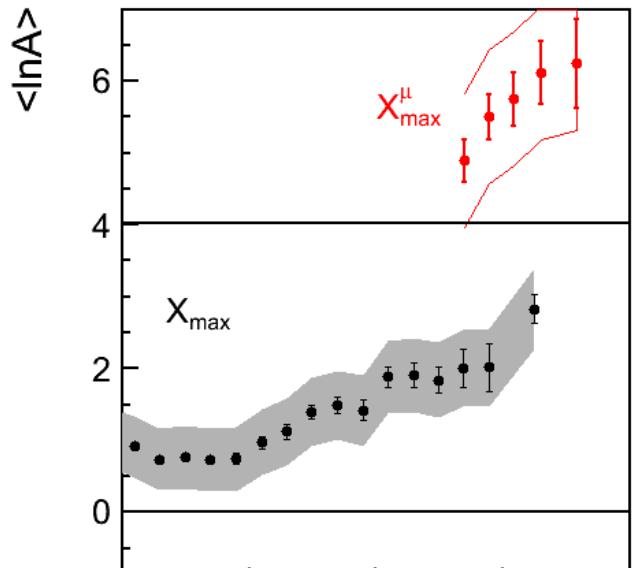
EM profile vs Muon Production profile



- ~1-2 hadr. generations after the 1st interaction
 - >90% energy is in the EM channel,
 - EM cascade practically decoupled from the hadronic cascade
- MPD profile: reflects development of the hadronic cascade
 - All hadr. generations contribute
- ΔX^μ & ΔX test different hadronic properties

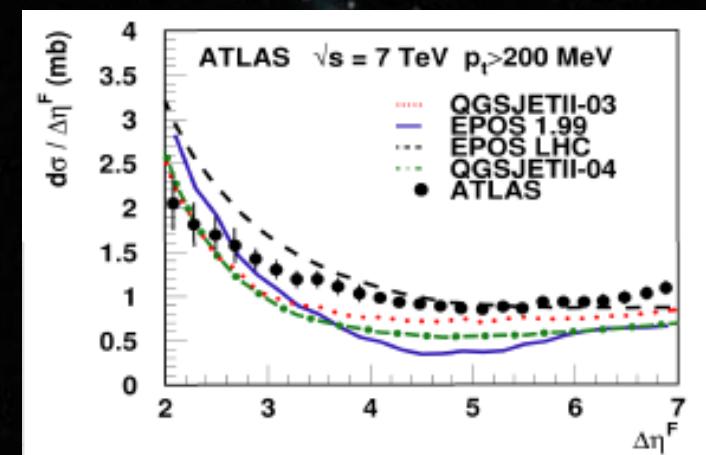
$$X_{\max} = X_1 + \Delta X$$

$$X_{\max}^\mu = X_1 + \Delta X^\mu$$



InA compatibility

Conversion of X_{max}^μ & X_{max} into InA using the models unveals problems (opportunities!) for the models.



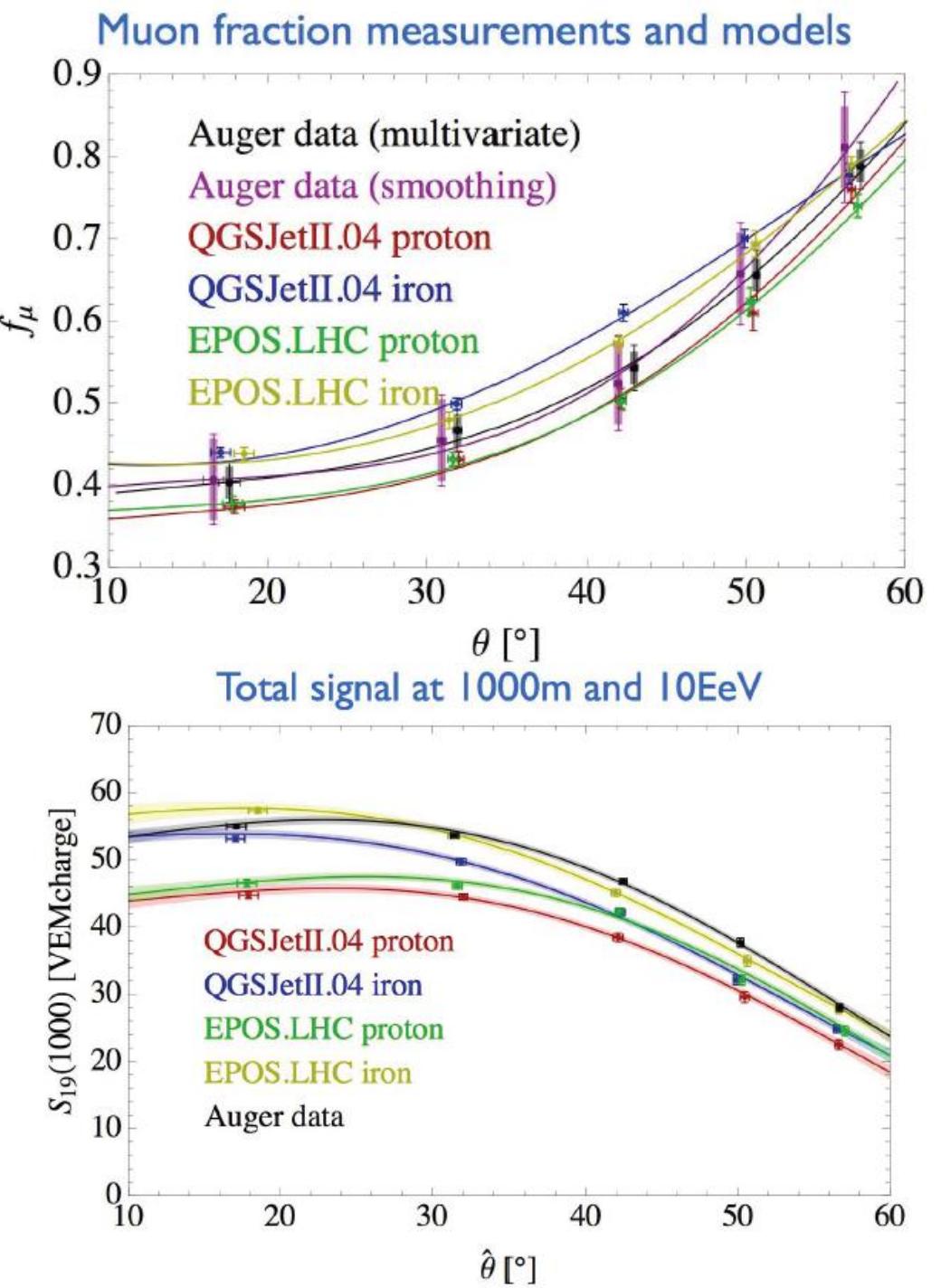
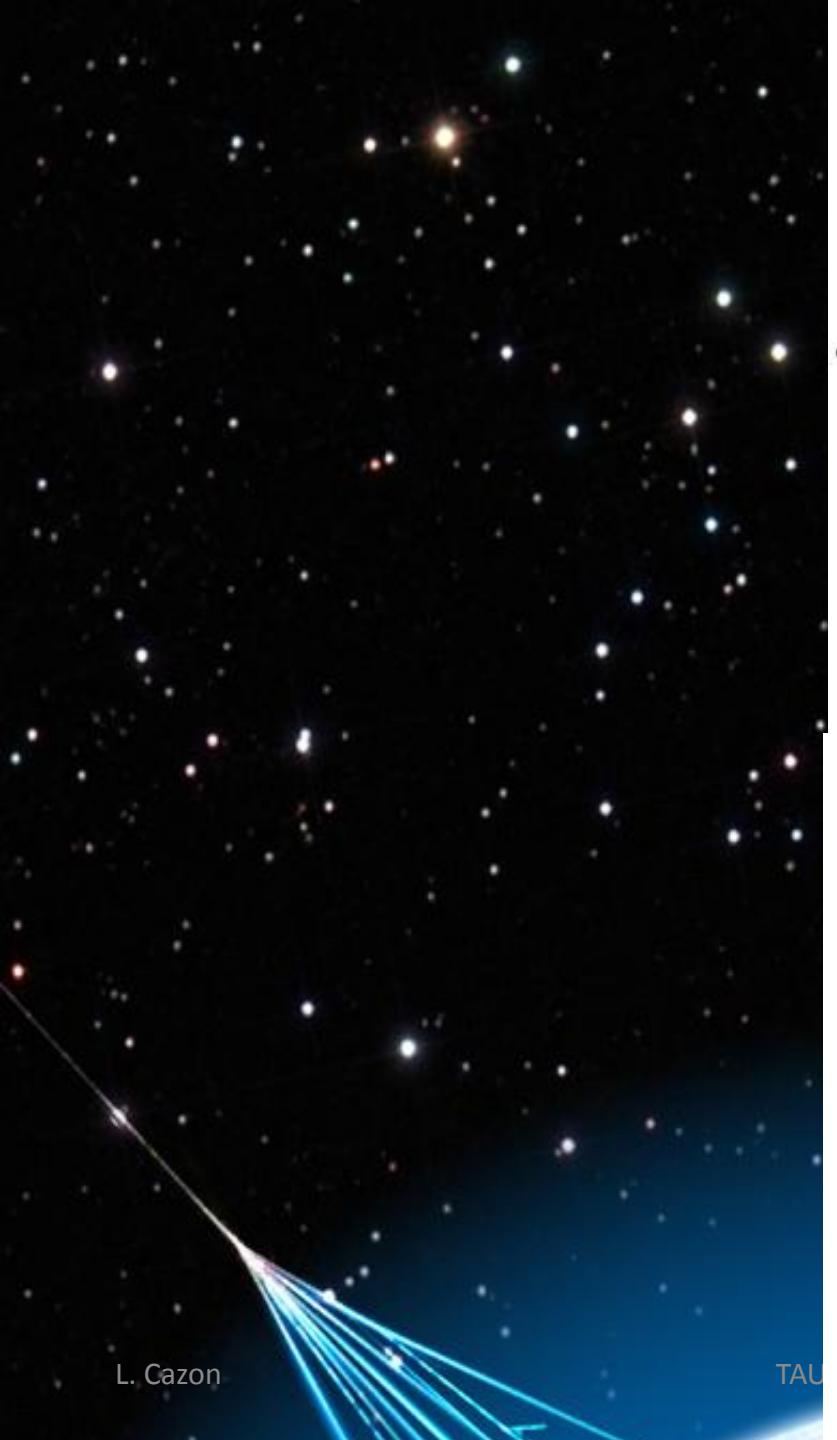
A difference between the models, LHC-EPOS makes a better treatment of diffraction rapidity gaps.

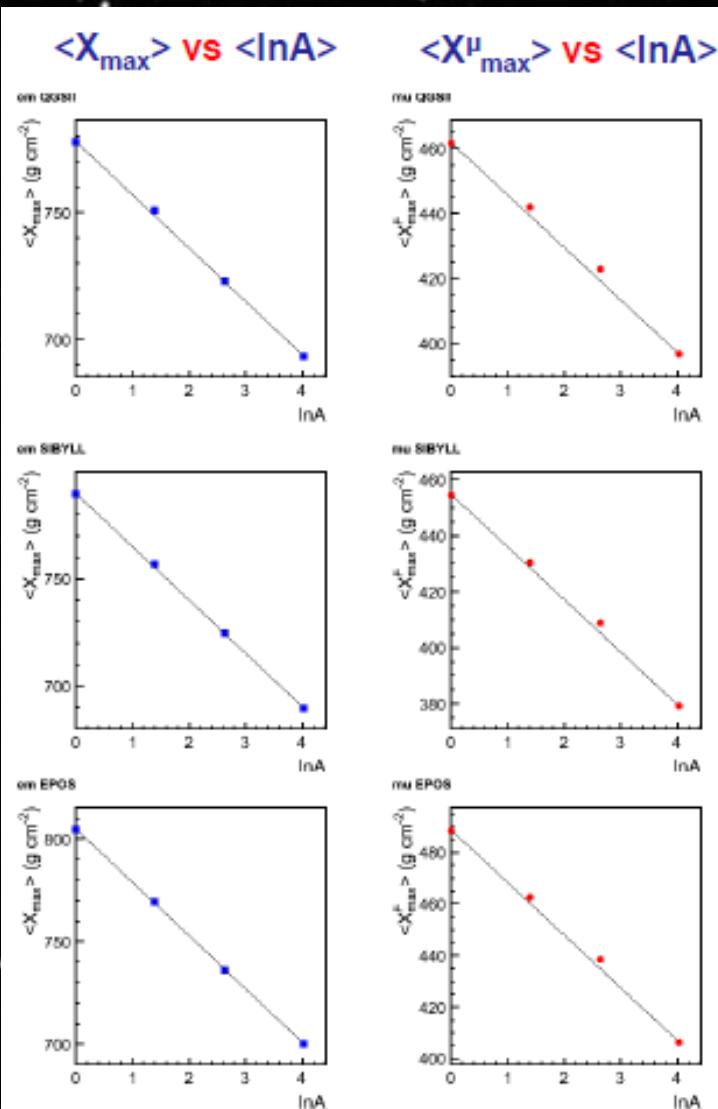
QGSJetII.04 could be missing additional and relevant effects that compensate X_{max}^μ

Conclusions

- $\sigma_{p\text{-Air}}$ for particle production measured at $\sqrt{s}=57 \text{ TeV}$
 - Compatible with most models.
 - Systematics: up to 25% He contamination.
- Updated measurements of muon production
 - new E scale, method's improvements and new hadronic models, **20% more muons.**
 - X_{\max} incompatible with iron dominated up to $2 \times 10^{19} \text{ eV} \rightarrow$
muon content is too large
 - Absolute value similar to iron predictions, but angular dependence close to proton at 10^{19} eV
 - Muon rescaling factor **1.3-1.6**
 - **No need for E rescaling**
- Muon Production Depth provides new constraints in hadronic models

BACK UP SLIDES





Linearity check*

$$\langle \ln A \rangle^{\text{EM}} = \ln 56 \frac{\langle X_{\max} \rangle_p - \langle X_{\max} \rangle}{\langle X_{\max} \rangle_p - \langle X_{\max} \rangle_{\text{Fe}}}$$

$$\langle \ln A \rangle^{\mu} = \ln 56 \frac{\langle X_{\mu_{\max}} \rangle_p - \langle X_{\mu_{\max}} \rangle}{\langle X_{\mu_{\max}} \rangle_p - \langle X_{\mu_{\max}} \rangle_{\text{Fe}}}$$

*Apparent MPD Indirectly Inferred using CONEX in combination with CORSIKA

4

